

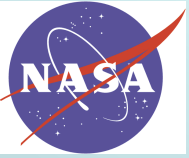
Kepler

A Search for Earth-size Planets

KEPLER Mission: Planets for Everyone



Steve B. Howell
& Kepler Team
AAVSO Fall 2009



Kepler Mission

Kepler

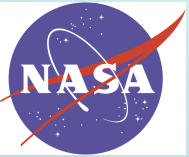
A Search for Earth-size Planets

**The Kepler space telescope's primary goal is to use transit photometry to detect exo-planets like Earth:
Rocky and in the Habitable Zone of a solar-like star.**

This mission is of great significance to our understanding of the universe and our place in it.

Kepler will obtain a statistically valid result by monitoring > 100,000 stars within a 100 sq. deg region of the sky.



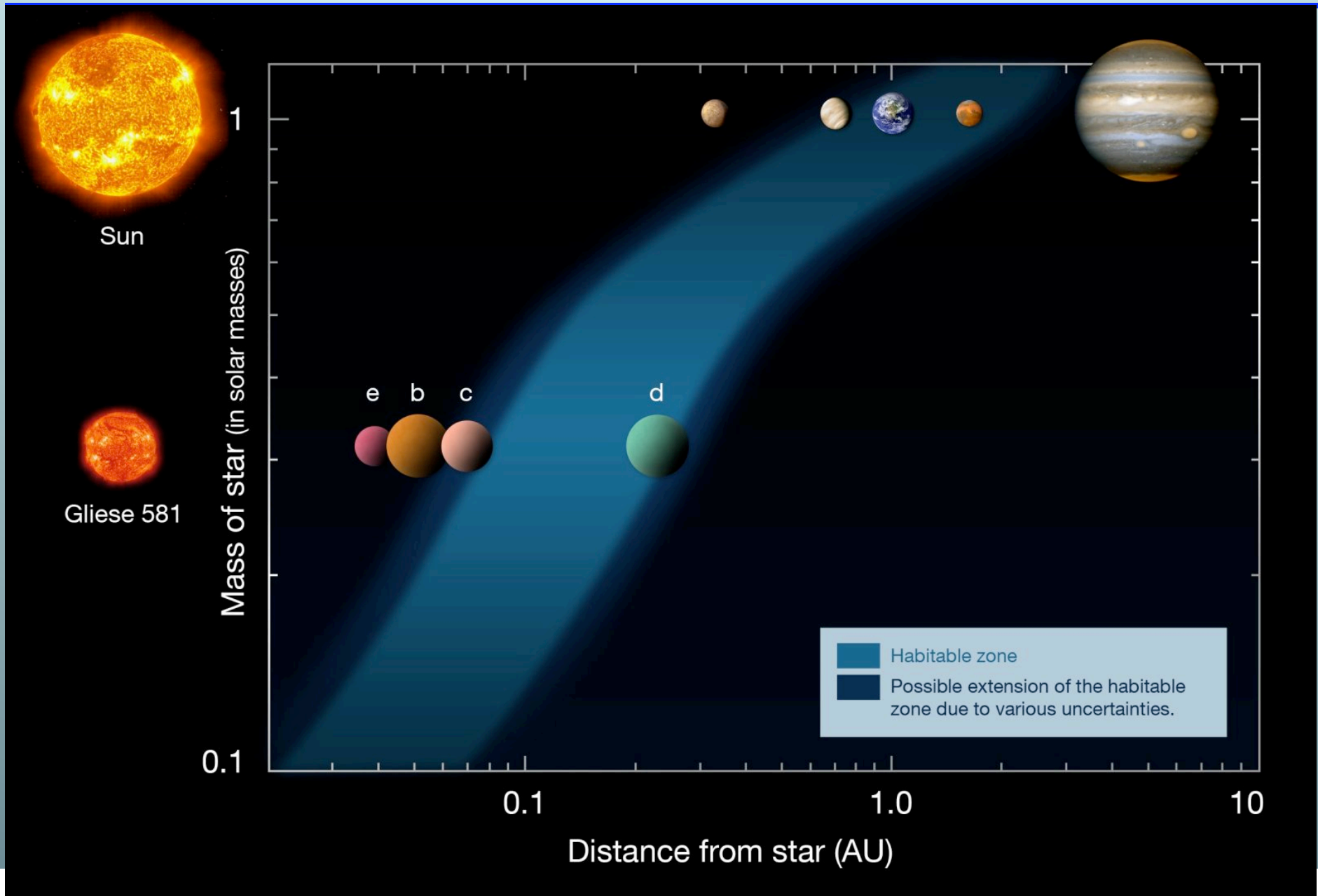


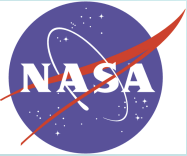
The Habitable Zone

May 19, 2019

Kepler

A Search for Earth-size Planets





EXTRA-SOLAR PLANETS – A Review



A Search for Earth-size Planets

What is an Exo-planet?

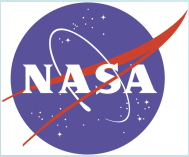
A planet that orbits a star other than our Sun.

The upper planet mass –the planet-star (brown dwarf) boundary - is fuzzy, but roughly a body orbiting a star with a mass ~ 13 - 15 Jupiter masses or less is considered a planet.

Late-type red dwarfs, brown dwarfs, and Jupiters are all \sim equal in radius

What is a Pirates Favorite Planet?

MARRRRRRRRRRRRRRRs!



TECHNIQUES FOR FINDING EXO-PLANETS

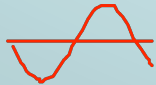


A Search for Earth-size Planets

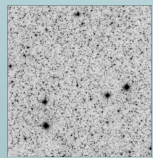
Method	Yield	Mass Limit	Status
--------	-------	------------	--------



Pulsar Timing	$m/M ; \tau$	Lunar	Successful
---------------	--------------	-------	------------



Radial Velocity	$m \sin i ; \tau$	Uranus	Successful
-----------------	-------------------	--------	------------



Astrometry	$m ; \tau ; D_s ; a$		
Ground: Telescope		Jupiter	Ongoing
Ground: Interferometer		<Jupiter	In development
Space: Interferometer		Uranus	Planned



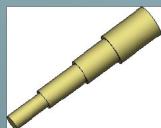
Transit Photometry	$A ; \tau ; \sin i = 1$		
Ground		Saturn	HD209458, OGLE TR-56?
Space		Earth	CoRoT, Kepler



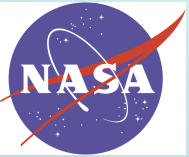
Reflection Photometry:	$albedo * A ; \tau$		
Space		Saturn	Planned Kepler



Microlensing:	$f(m, M, r, D_s, D_L)$		
Ground		sub-Uranus	On-going



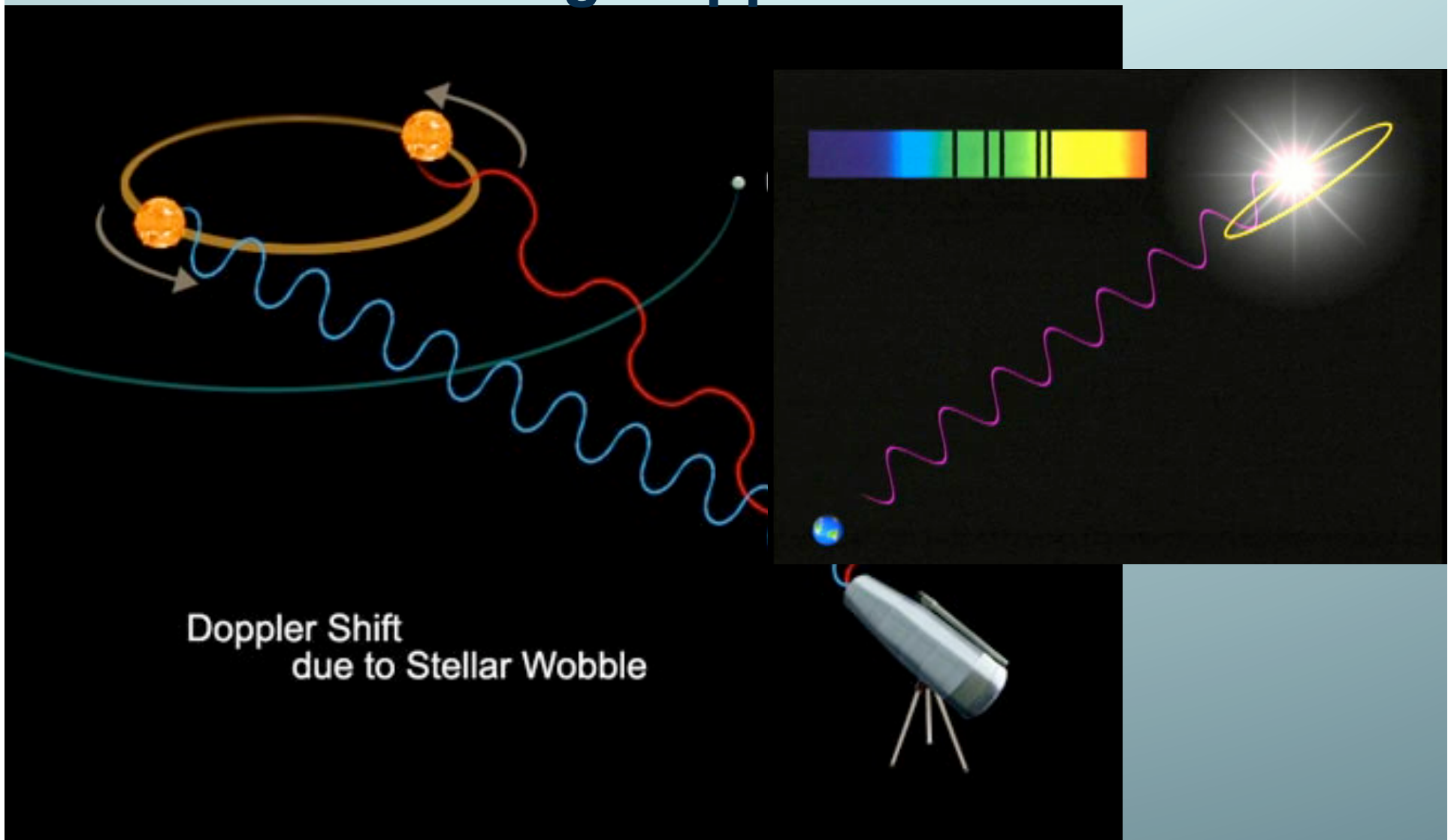
Direct Imaging	$albedo * A ; \tau ; D_s ; a ; M$		
Ground		Saturn	Being studied
Space		Earth	Being studied?



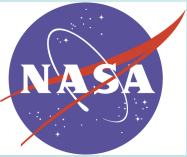
How to Detect Planets using Doppler Shift

Kepler

A Search for Earth-size Planets



Doppler Shift
due to Stellar Wobble

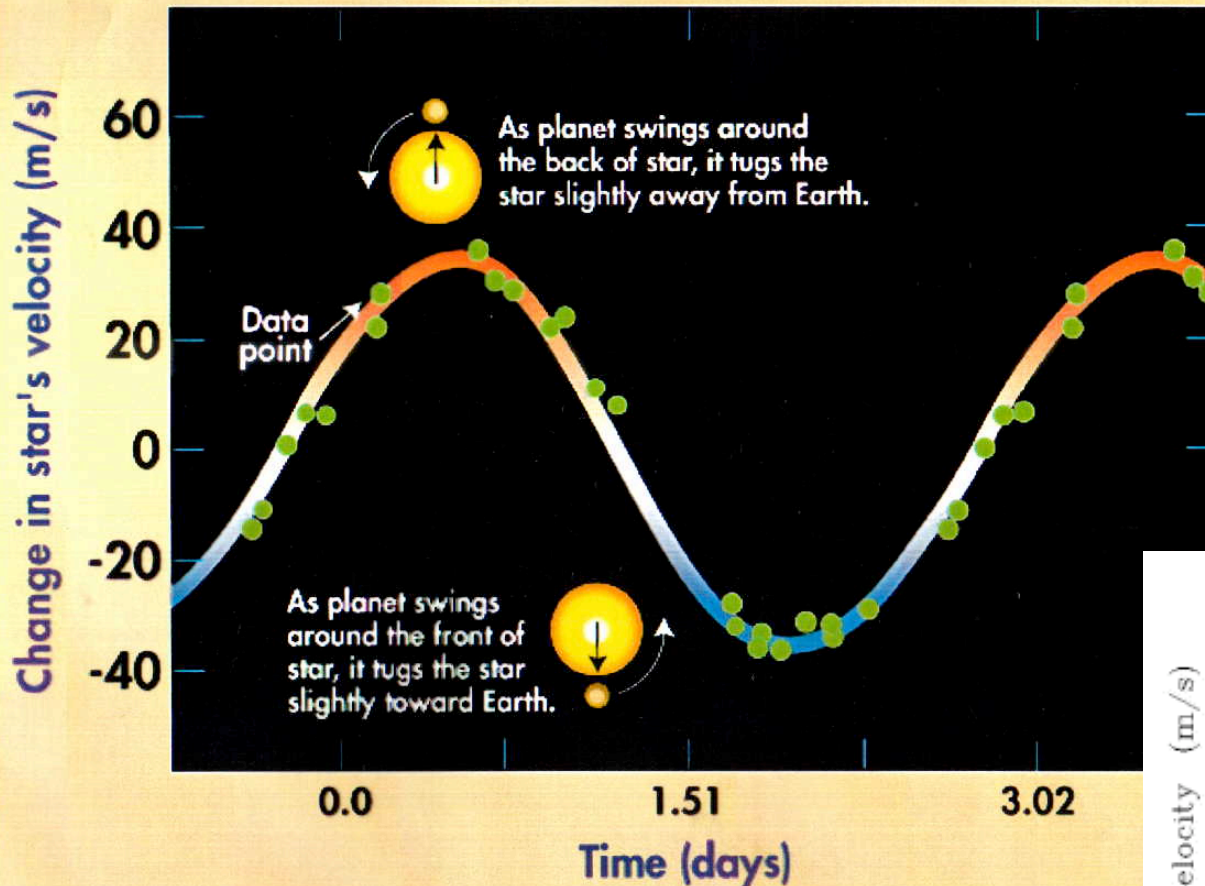


Doppler shift planet discovery

Kepler

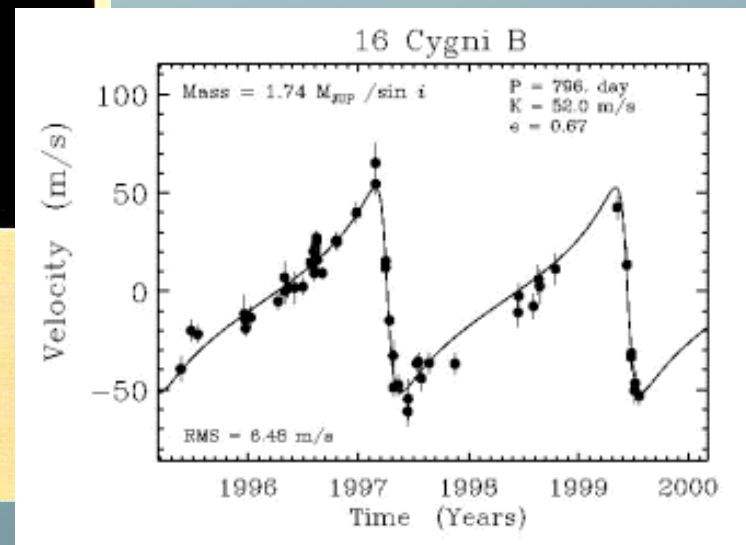
A Search for Earth-size Planets

Planet Orbiting Star HD46375



Orbital period = 3.02 days
Orbiting planet = 80% of Saturn's mass

The “wobble” method gets the orbital period, semi-major axis, and a lower limit on the mass of the planet. RVs have detected down to 7 Earth-mass planets very close in, (but favors gas giant planets).





Direct Imaging !

Kepler

A Search for Earth-size Planets

Planets Orbiting HR 8799 (Sept. 2008)

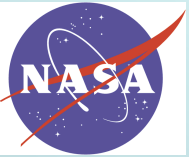


July 2004
+

July 2004
+

July 2008
+

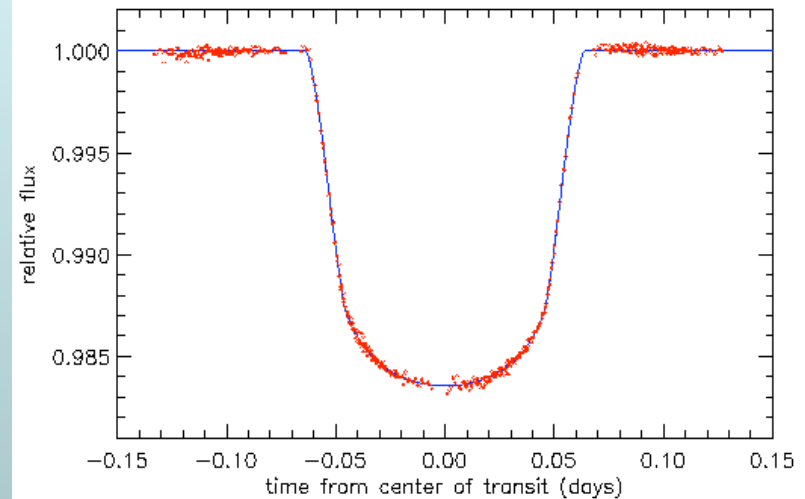
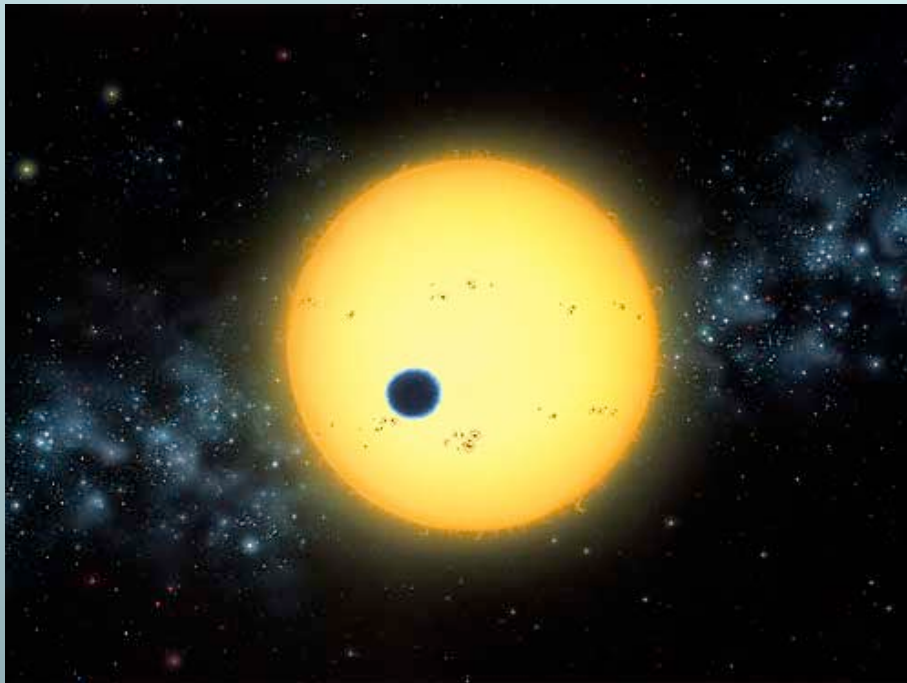
0.5 arcsec
20 AU



Kepler will use the Transit Method

Kepler

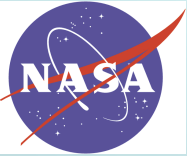
A Search for Earth-size Planets



HST measurement of HD209458

Transits tell us: Planet radius, orbit size, & density

Likelihood of transit is governed by orbital plane geometry. About 10% for very close-in large planets while only $\sim 0.5\%$ for true-Earth analogs. Short orbital periods and larger planets are favored.



CURRENT EXO-PLANET SURVEYS



A Search for Earth-size Planets

- **82 Ground-based exo-planet surveys are underway**
 - Doppler, transit, imaging, astrometry
- **21 space missions: 5 are current, remaining planned**
 - Current dedicated missions use transits

- **Dedicated Planet Missions w/ ability to find Earths**
CoRoT (CNES)

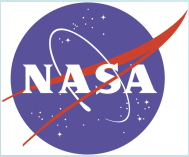
Kepler (NASA)

- **Useful Planet Finders/Follow-up Missions**

MOST (Canada)

HST (photometry, astrometry)

Spitzer Space telescope (thermal imaging)



EXTRA-SOLAR PLANETS



A Search for Earth-size Planets

What have we found so far? (Pre-Kepler, November 1, 2009)

Radial velocity or astrometry

320 planetary systems 376 planets 38 multiple planets

Transiting planets

62 planetary systems 62 planets 3 multiple planets

Candidates detected by microlensing

8 planetary systems 9 planets 1 multiple planet

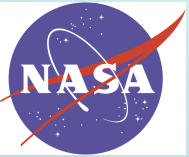
Candidates detected by imaging

9 planetary systems 11 planets 1 multiple planet

Candidates detected by timing

4 planetary systems 7 planets 2 multiple planets

403 Total Exo-planets



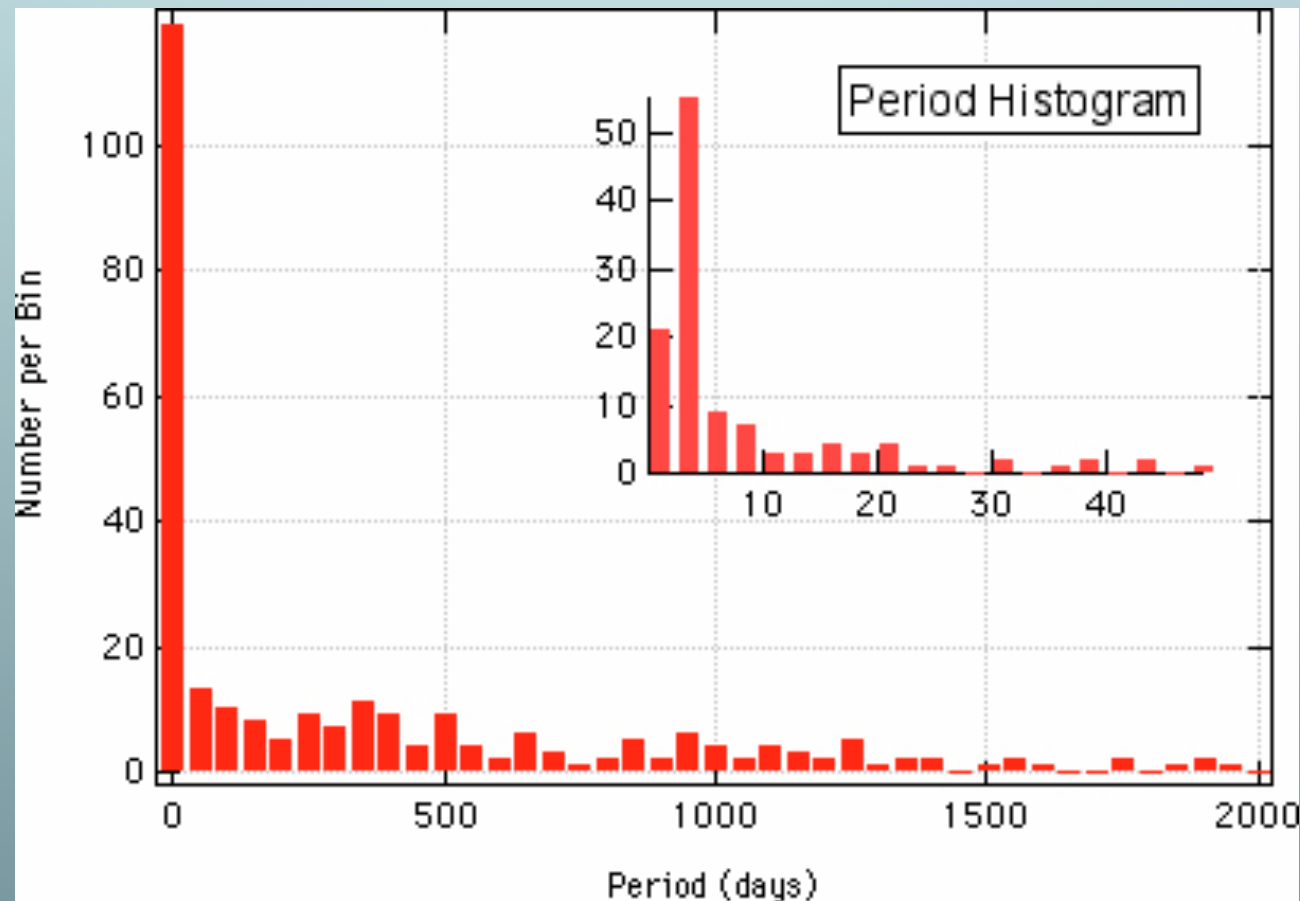
EXTRA-SOLAR PLANETS

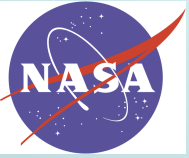


A Search for Earth-size Planets

Characteristics:

Orbital Period (days) vs. Number of known Exo-Planets





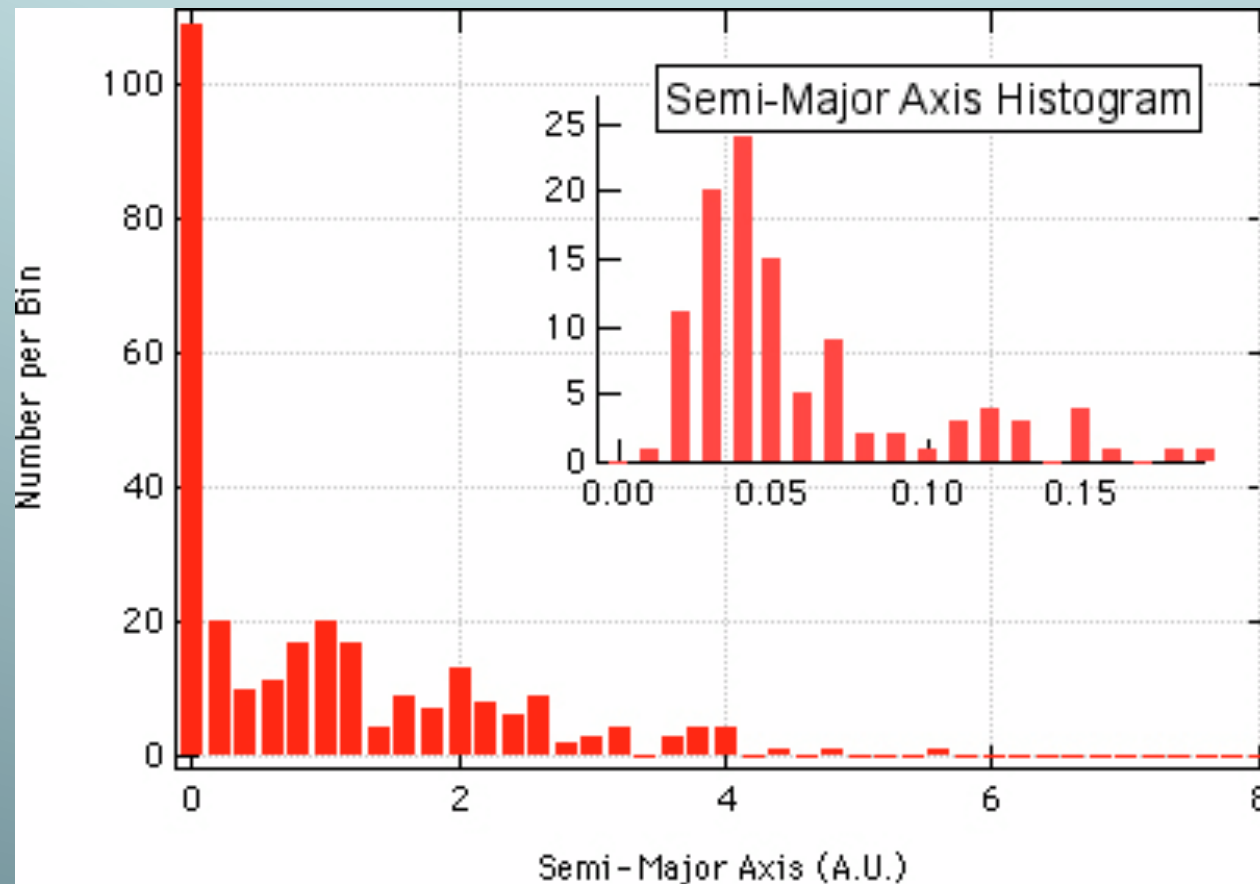
EXTRA-SOLAR PLANETS

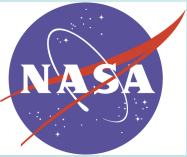


A Search for Earth-size Planets

Characteristics:

Semi-Major Axis (AU) vs. Number of known Exo-Planets





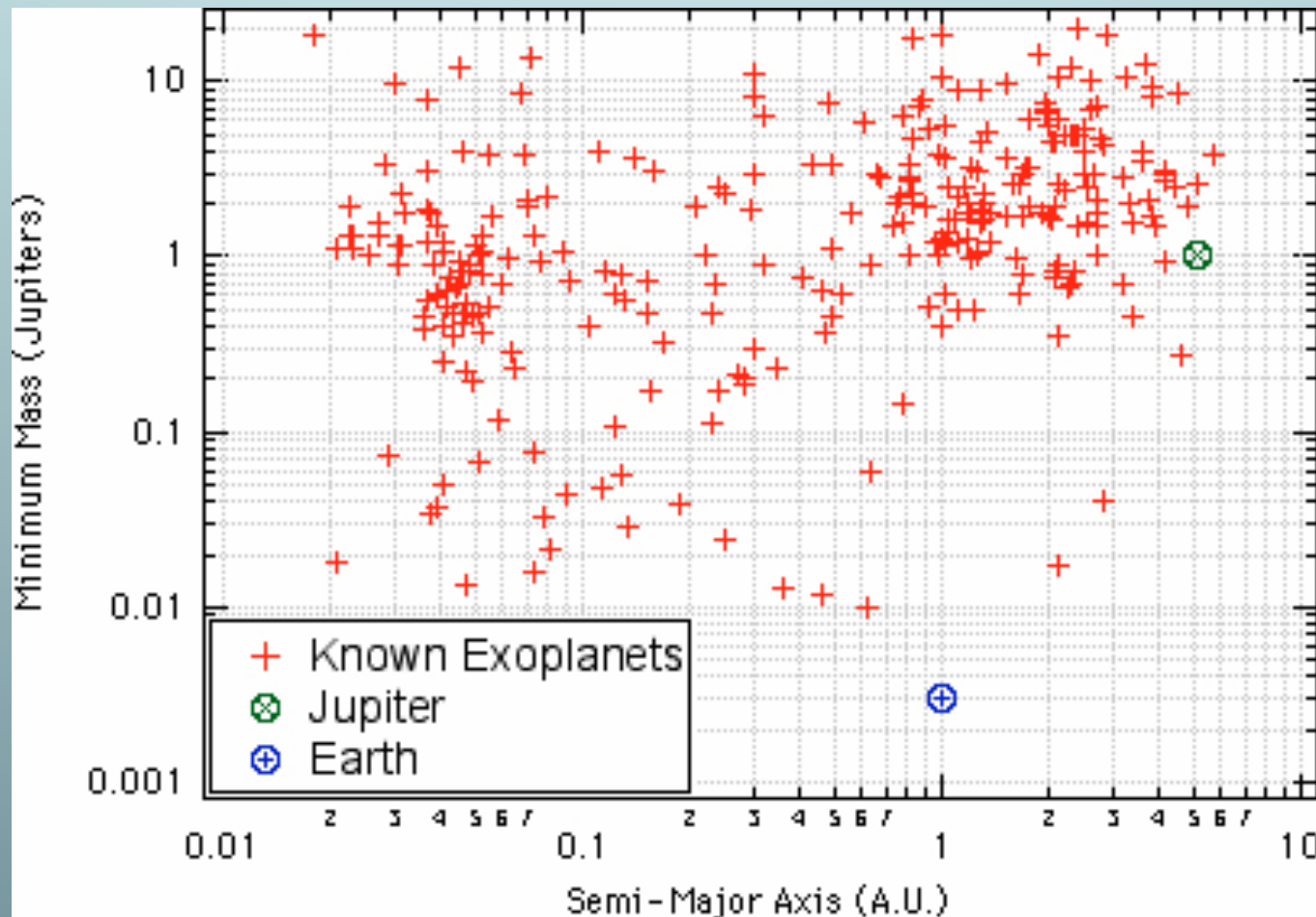
EXTRA-SOLAR PLANETS

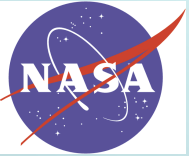


A Search for Earth-size Planets

Characteristics:

Mass vs. Semi-Major axis



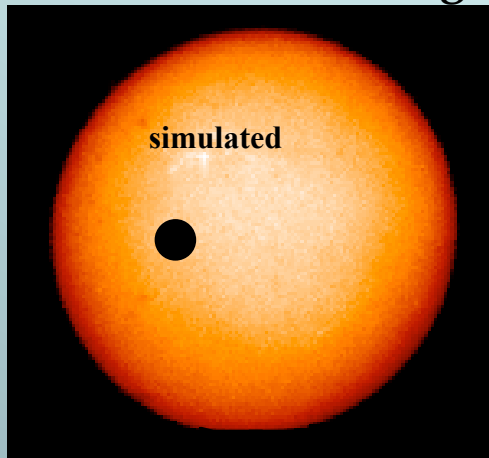


How to Detect Earth-Size Planets

Kepler

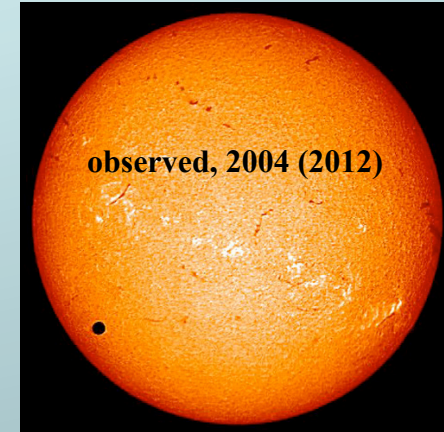
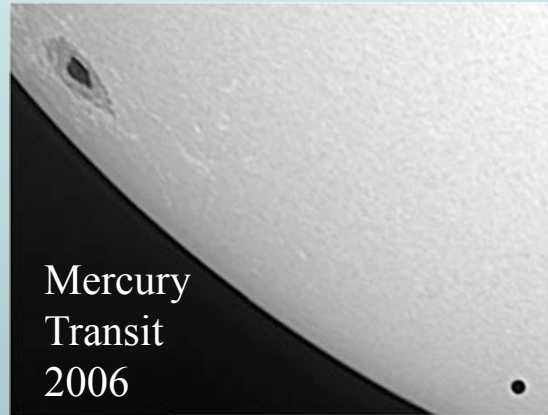
A Search for Earth-size Planets

- The relative change in brightness is equal to the area ratio: $A_{\text{planet}}/A_{\text{star}}$



Jupiter:

1% area of the Sun (1/100)



Earth or Venus

0.01% area of the Sun (1/10,000)

- To measure 0.01% must get above the Earth's atmosphere
- Allows continuous observation (no pesky Sun getting in the way, no Weather issues)
- Patience required:
Must observe 3+ transits, with same brightness drop, duration and period: near 3 years to complete



KEPLER MISSION WILL ADDRESS

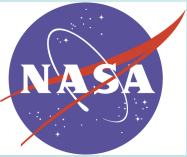
Kepler

A Search for Earth-size Planets

- **What is the frequency of Earth-size planets in or near the Habitable Zone (HZ) of solar-like stars?**
- **What are the frequency & orbital distributions of planets in star systems?**



- **What are the distributions of semi-major axes, albedo, size, and mass, of short-period planets?**



Kepler Mission

Kepler

A Search for Earth-size Planets

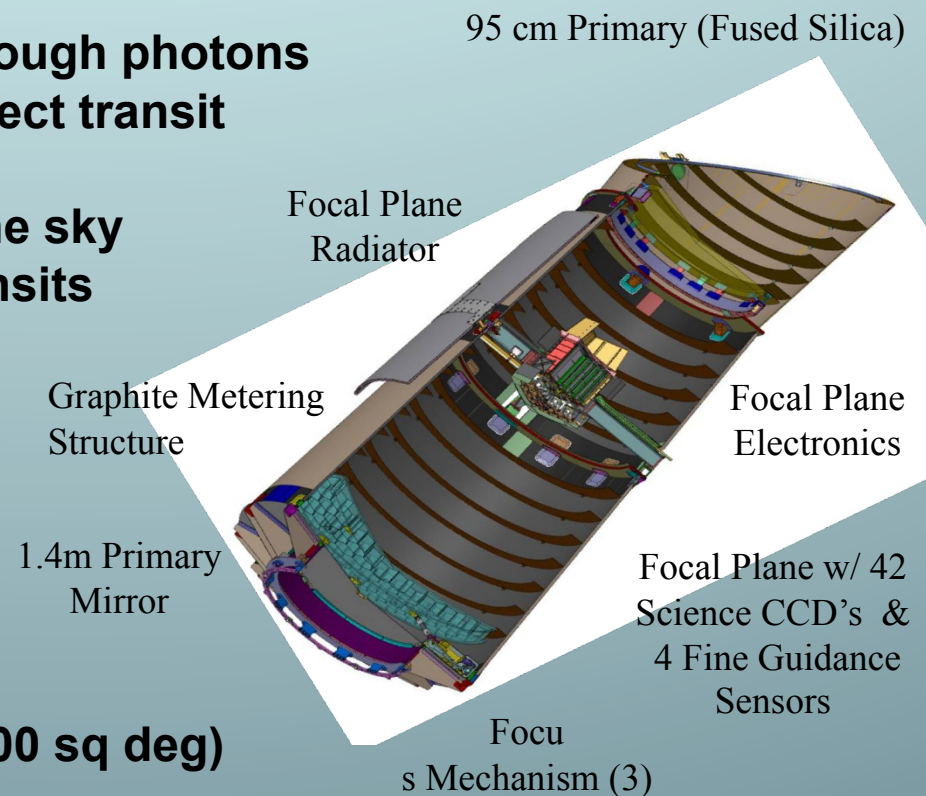
KEPLER: A Wide Field-of-View Photometer that Monitors 100,000+ Stars for 3.5 yrs with Enough Precision to Find Earth-size Planets in the Habitable Zone

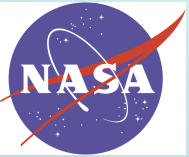
Use transit photometry to detect Earth-size planets

- **0.95 meter aperture provides enough photons**
- **Observe for several years to detect transit patterns**
- **Monitor a single large area on the sky continuously to avoid missing transits**
- **Use heliocentric orbit**
- **Up to 170,000 targets at 30 min cadence & 512 at 1 min**

Get statistically valid results by monitoring; 100,000 stars

- **Wide Field-of-view telescope (100 sq deg)**
- **Large array of CCD detectors**





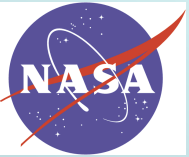
PREPARING FOR LAUNCH

Kepler

A Search for Earth-size Planets

- CCD Focal Plane being assembled





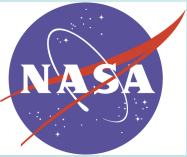
PREPARING FOR LAUNCH

Kepler

A Search for Earth-size Planets

- Spacecraft bus & Primary Mirror

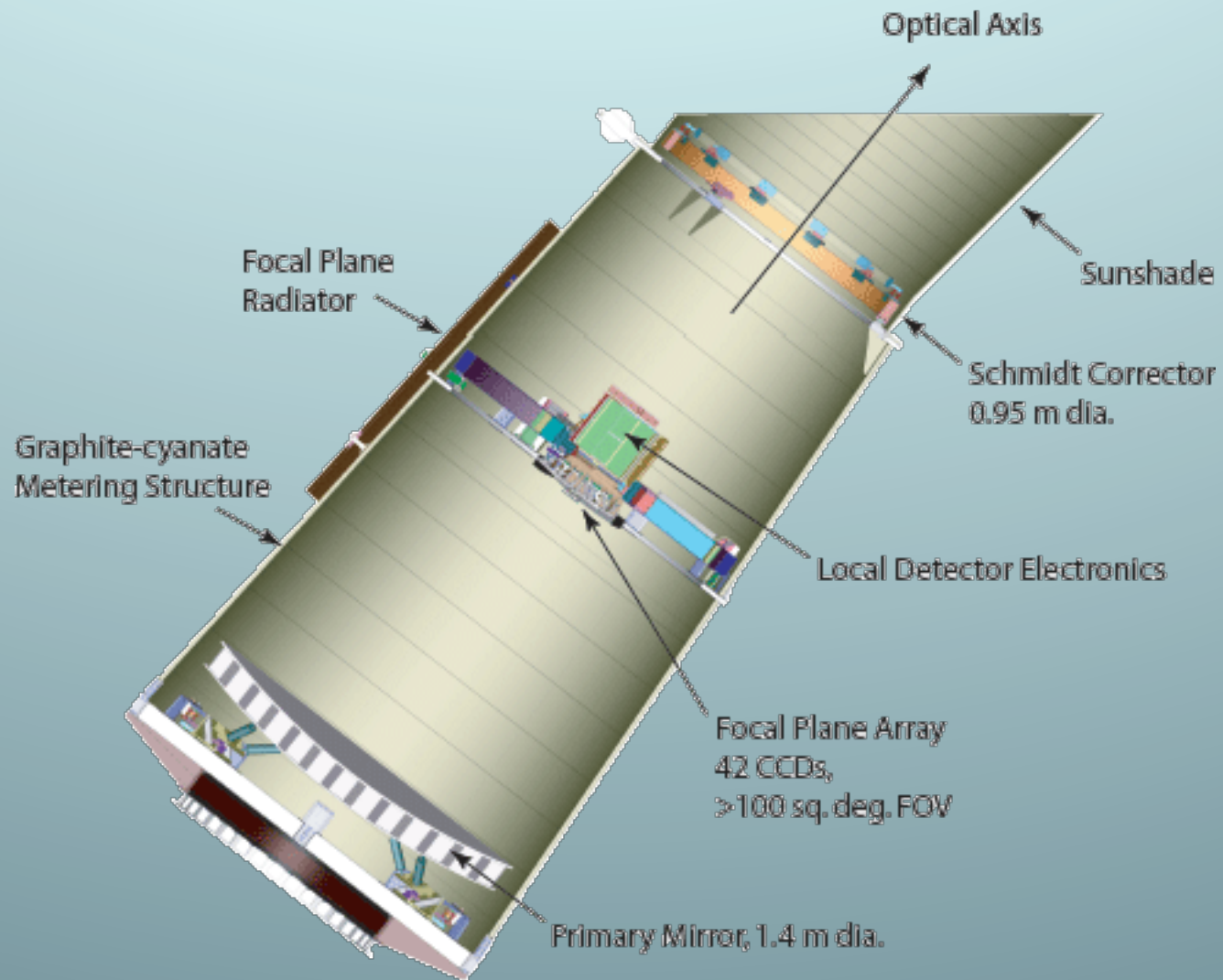


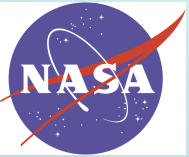


Telescope Schematic

Kepler

A Search for Earth-size Planets

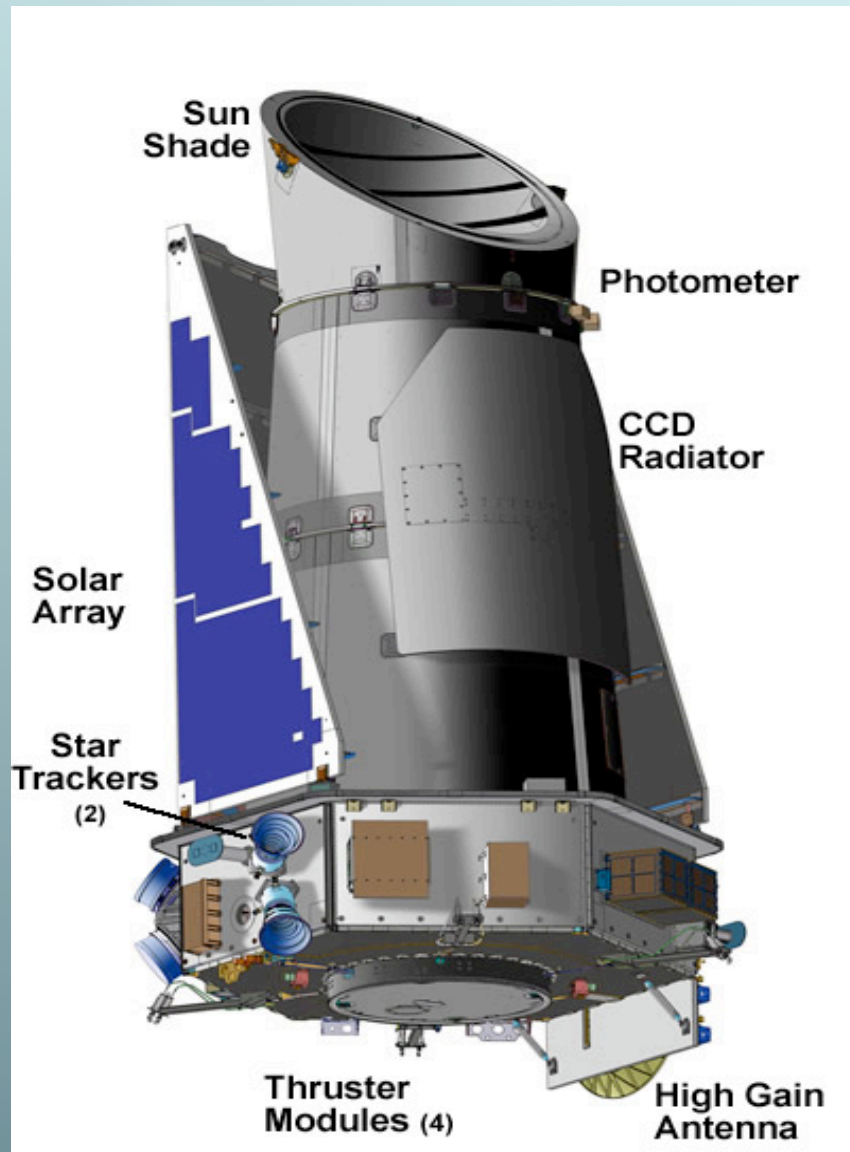




Drawing and Reality

Kepler

A Search for Earth-size Planets



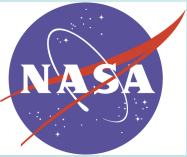


LAUNCH ON MARCH 6, 2009

Kepler

A Search for Earth-size Planets

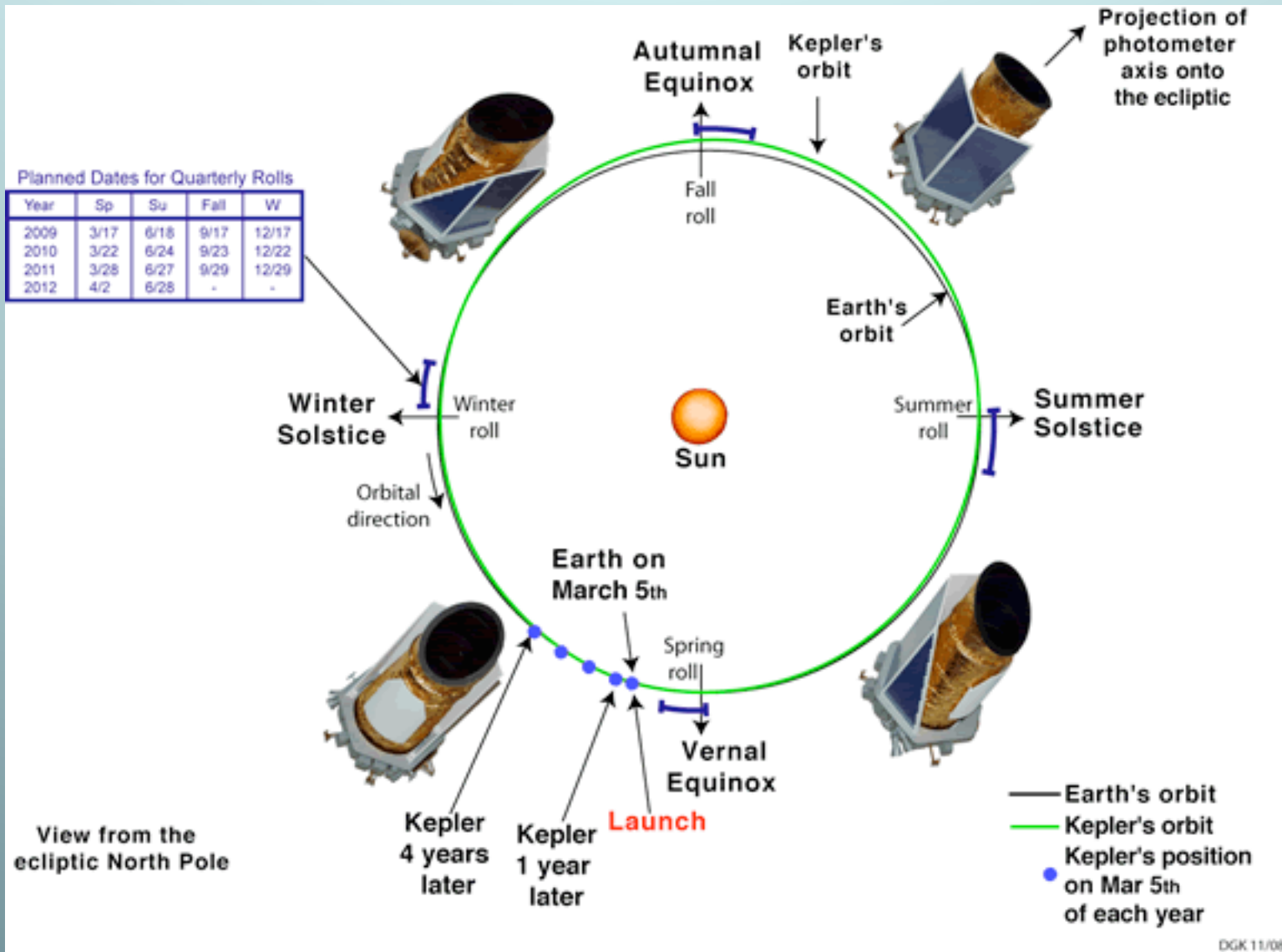


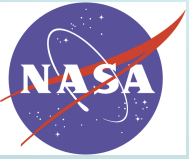


KEPLER SPACECRAFT ORBIT

Kepler

A Search for Earth-size Planets

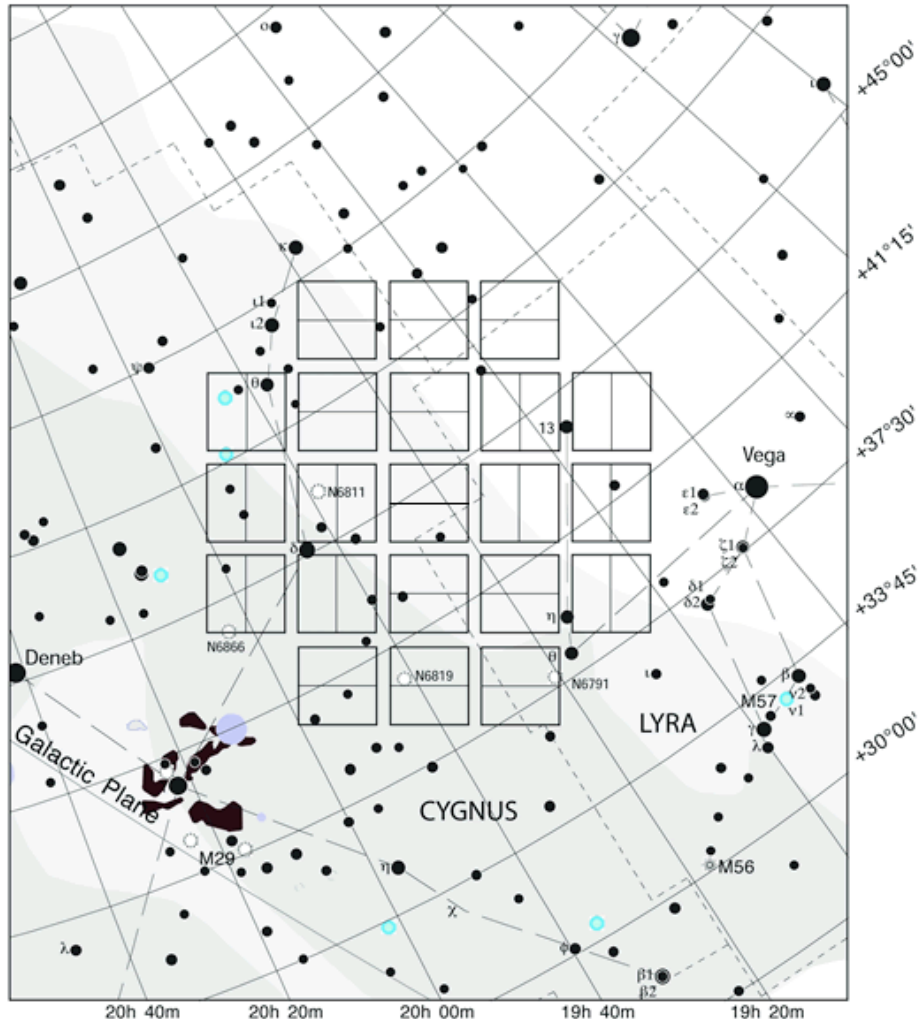




Kepler Field of View

Kepler

A Search for Earth-size Planets



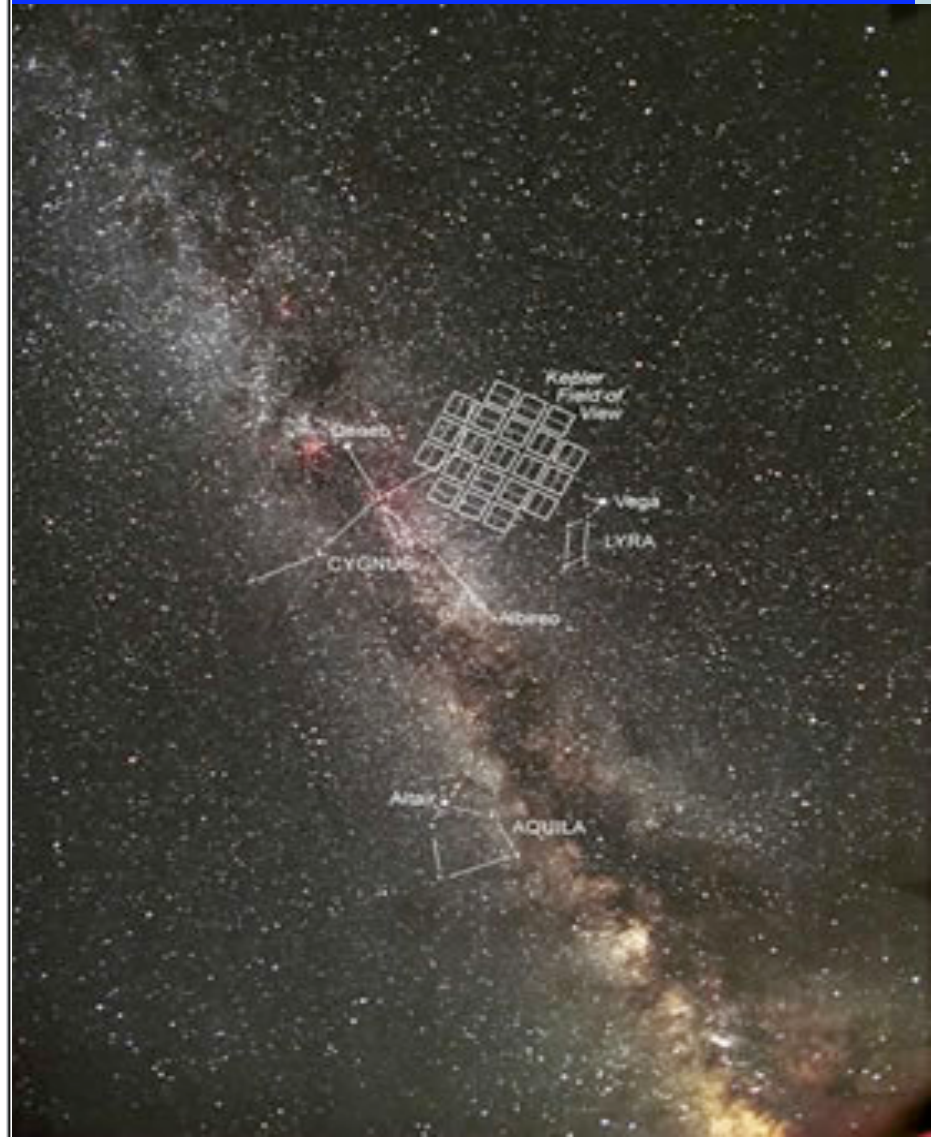
Star Magnitudes
0 1 2 3 4 5 6

Kepler FOV

○ Open Cluster
⊙ Globular Cluster
☁ Nebula
⊕ Planetary Nebula

THE SKY
Astronomy Software

FOV Center RA: 19h 22m 40s Dec: +44 30' 00" 9/10/04

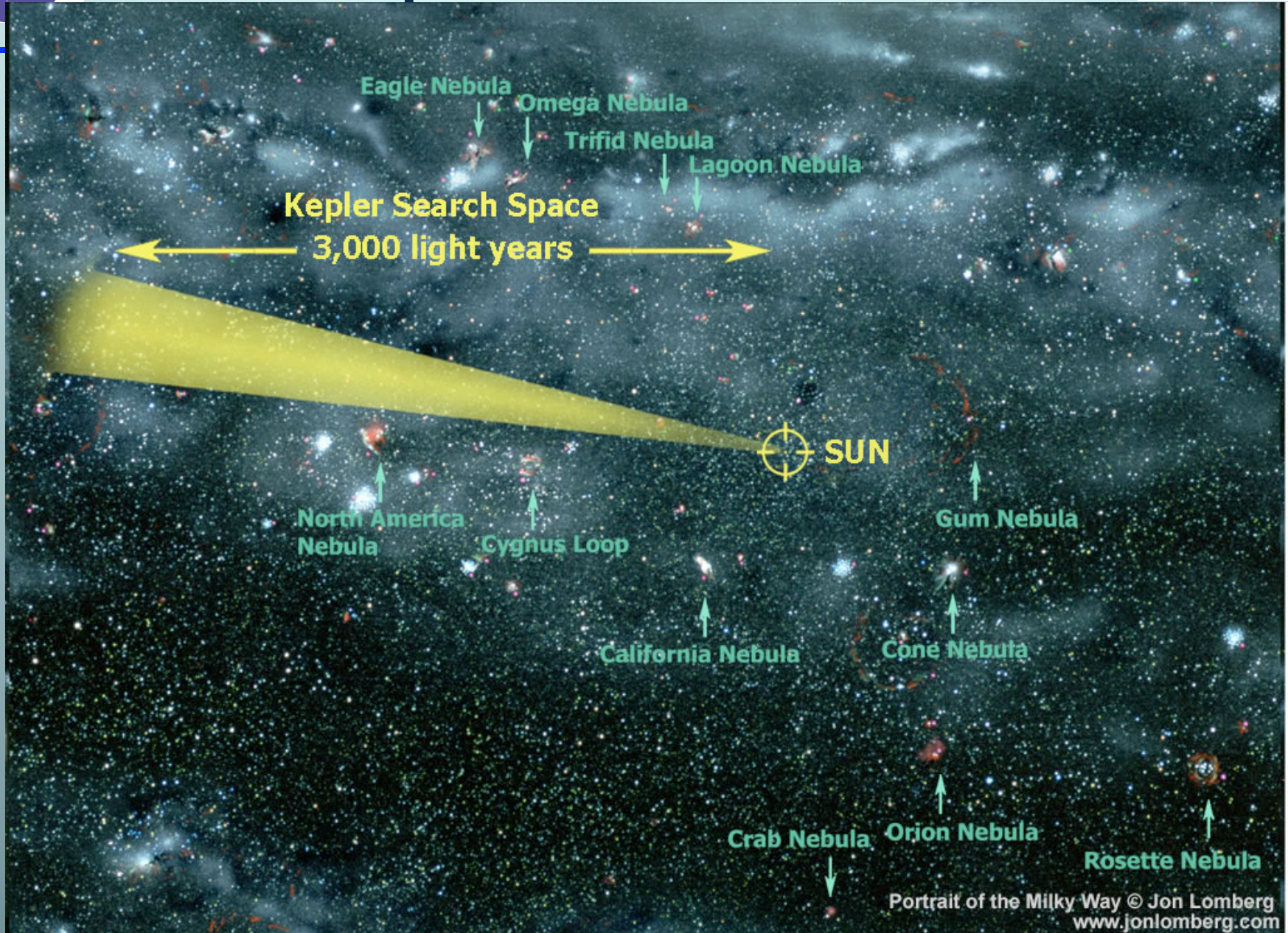




Kepler Field of View

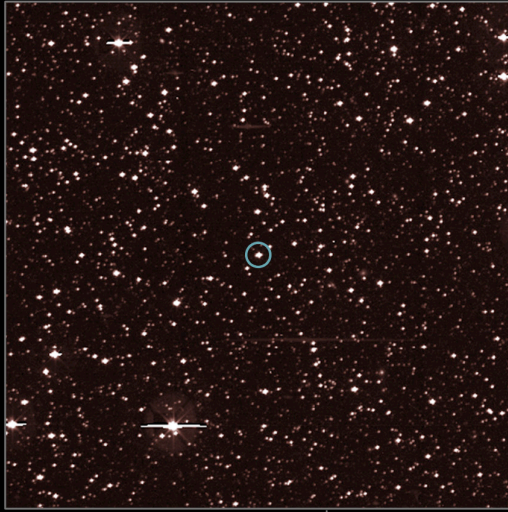
Kepler

Planets

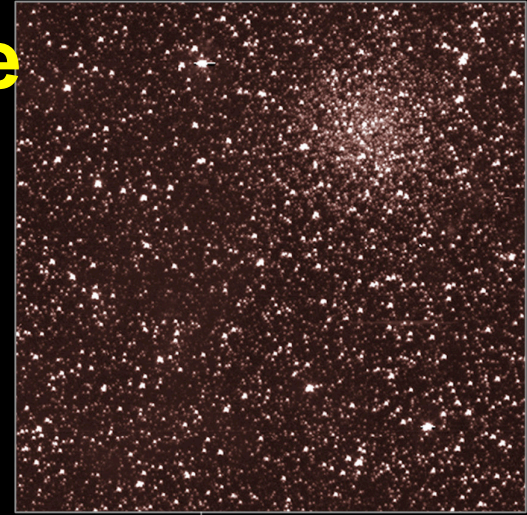


Portrait of the Milky Way © Jon Lomberg
www.jonlomberg.com

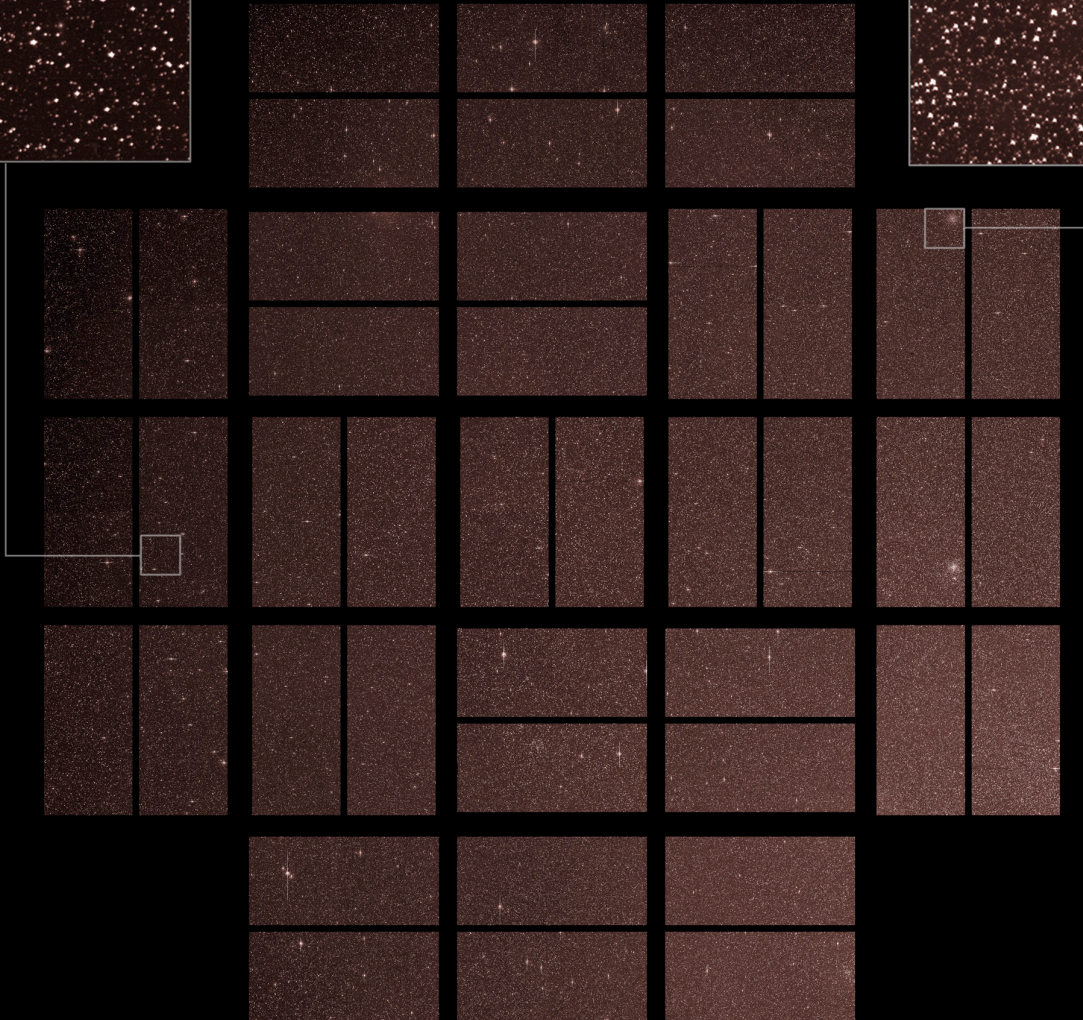
First Light Image

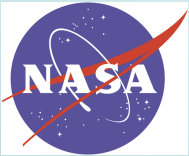


TrES-2



NGC 6791

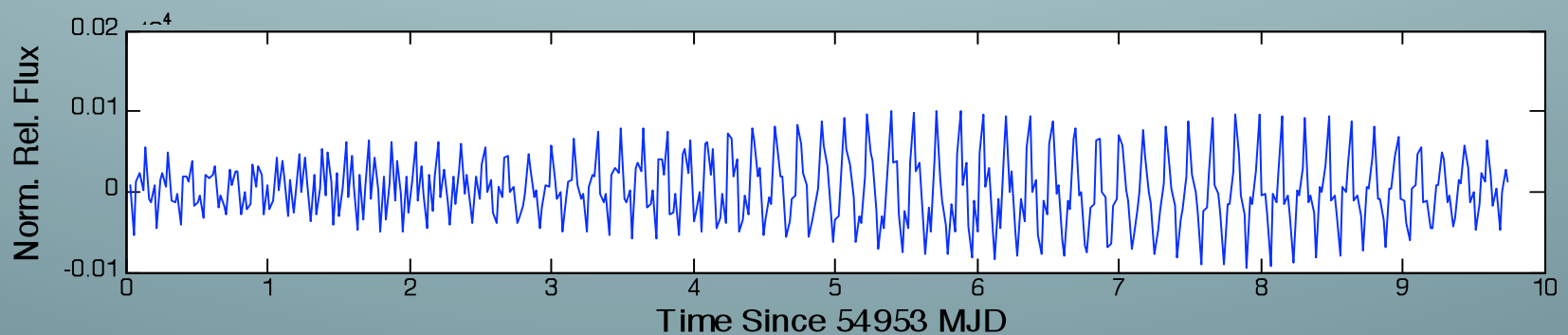
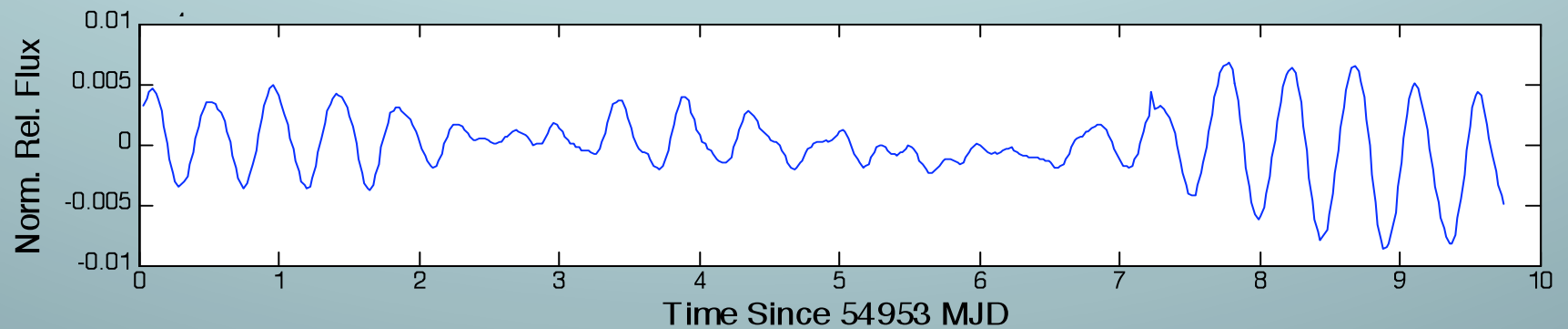
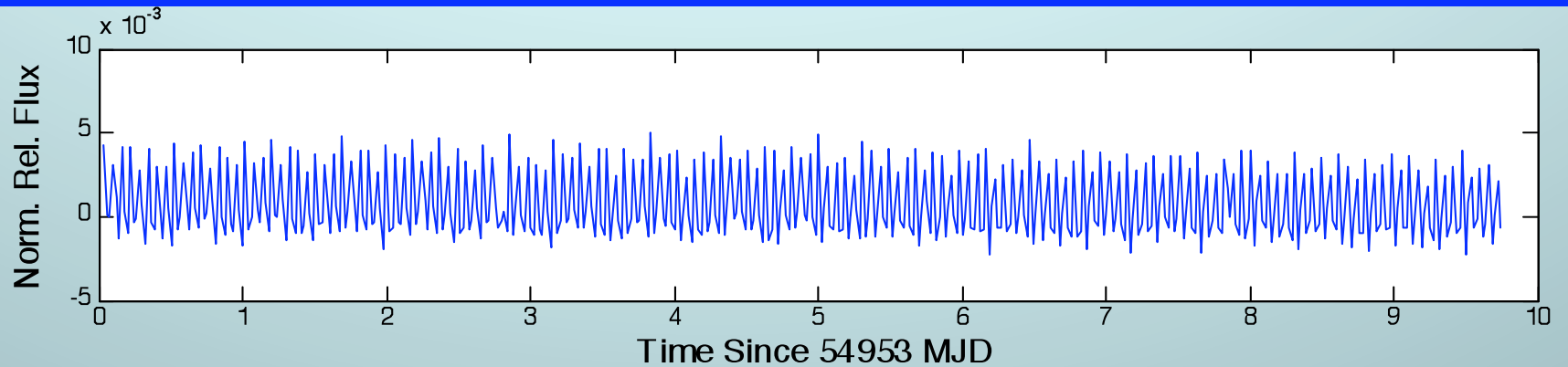


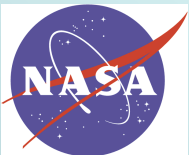


SAMPLE LIGHT CURVES



A Search for Earth-size Planets

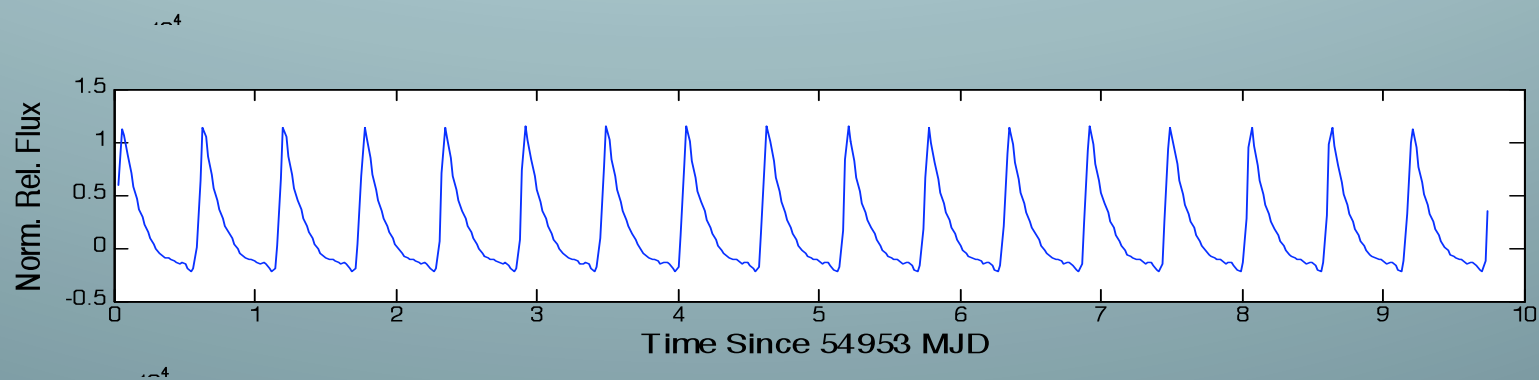
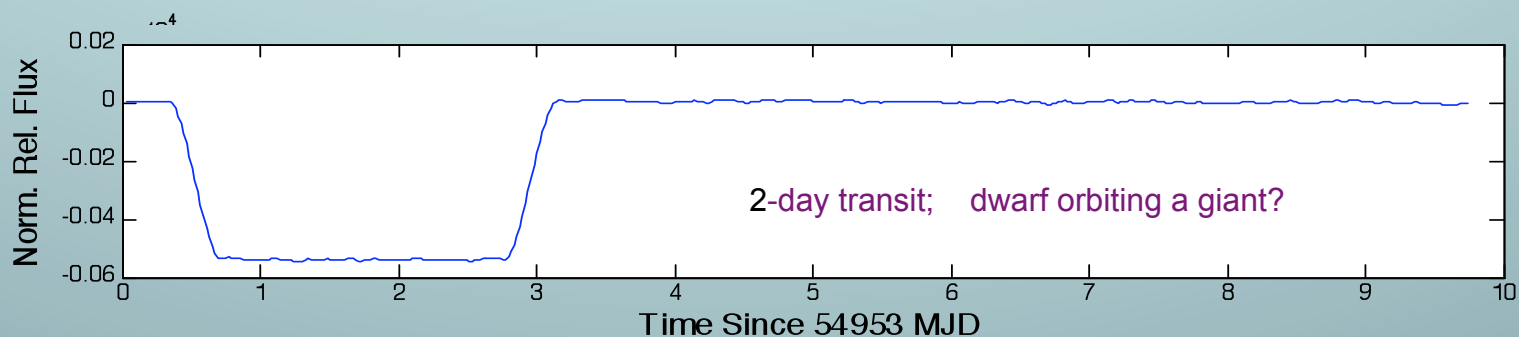
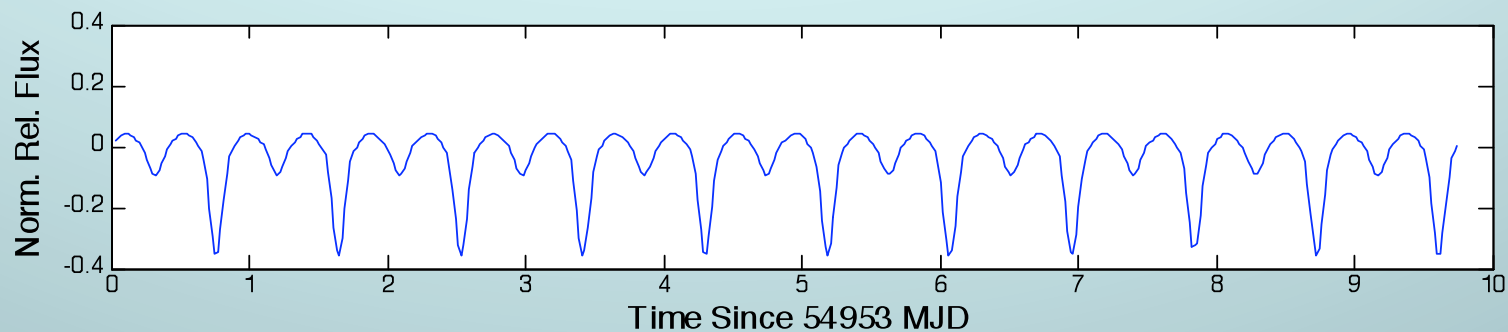


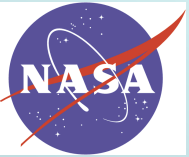


SAMPLE LIGHT CURVES



A Search for Earth-size Planets





MEASURED LIGHT CURVE NOISE, SO FAR

Kepler

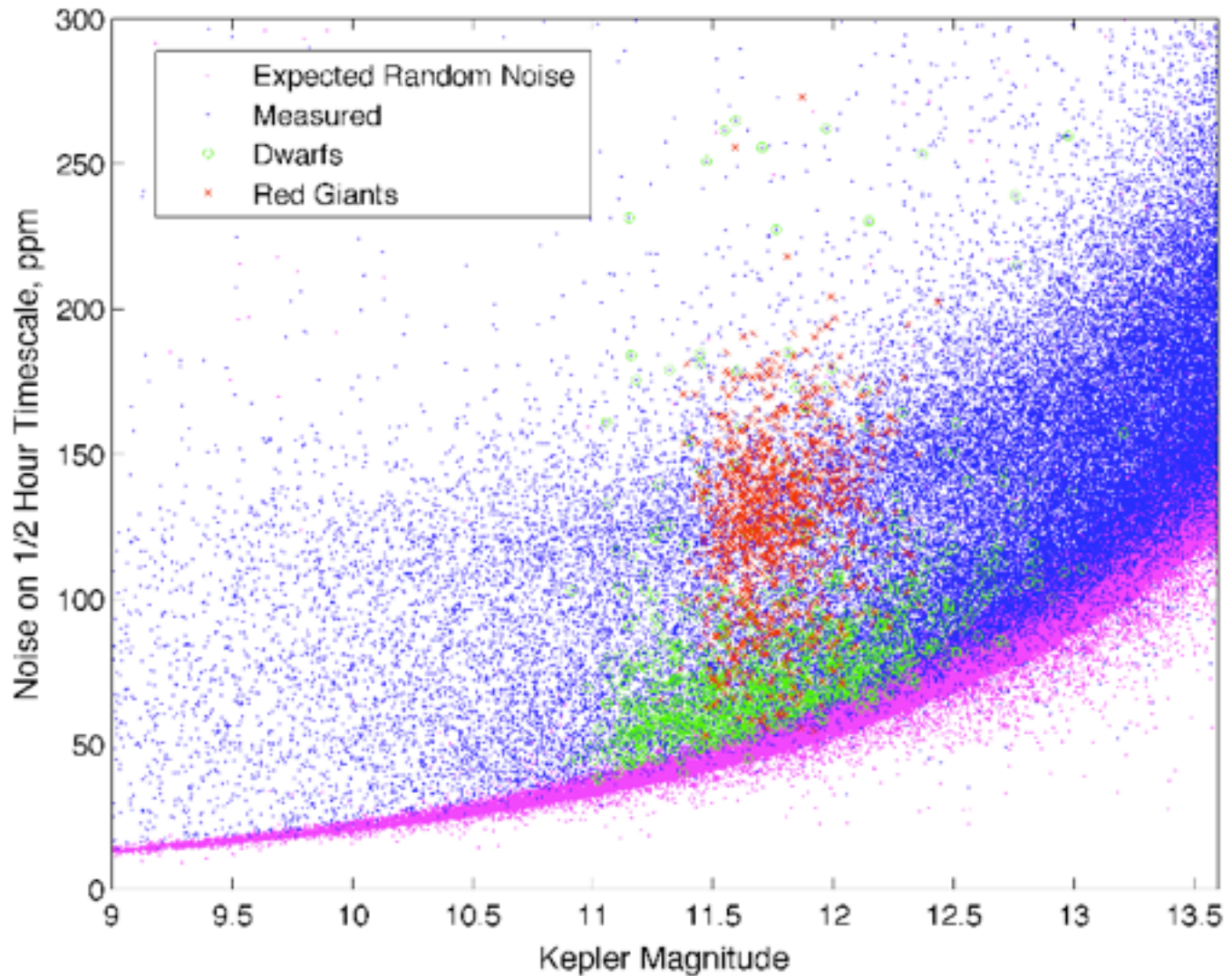
A Search for Earth-size Planets

Magenta =
expected
noise (non-
variable)

Blue =
measured

Green =
1000 dwarfs

Red = 1000
giants

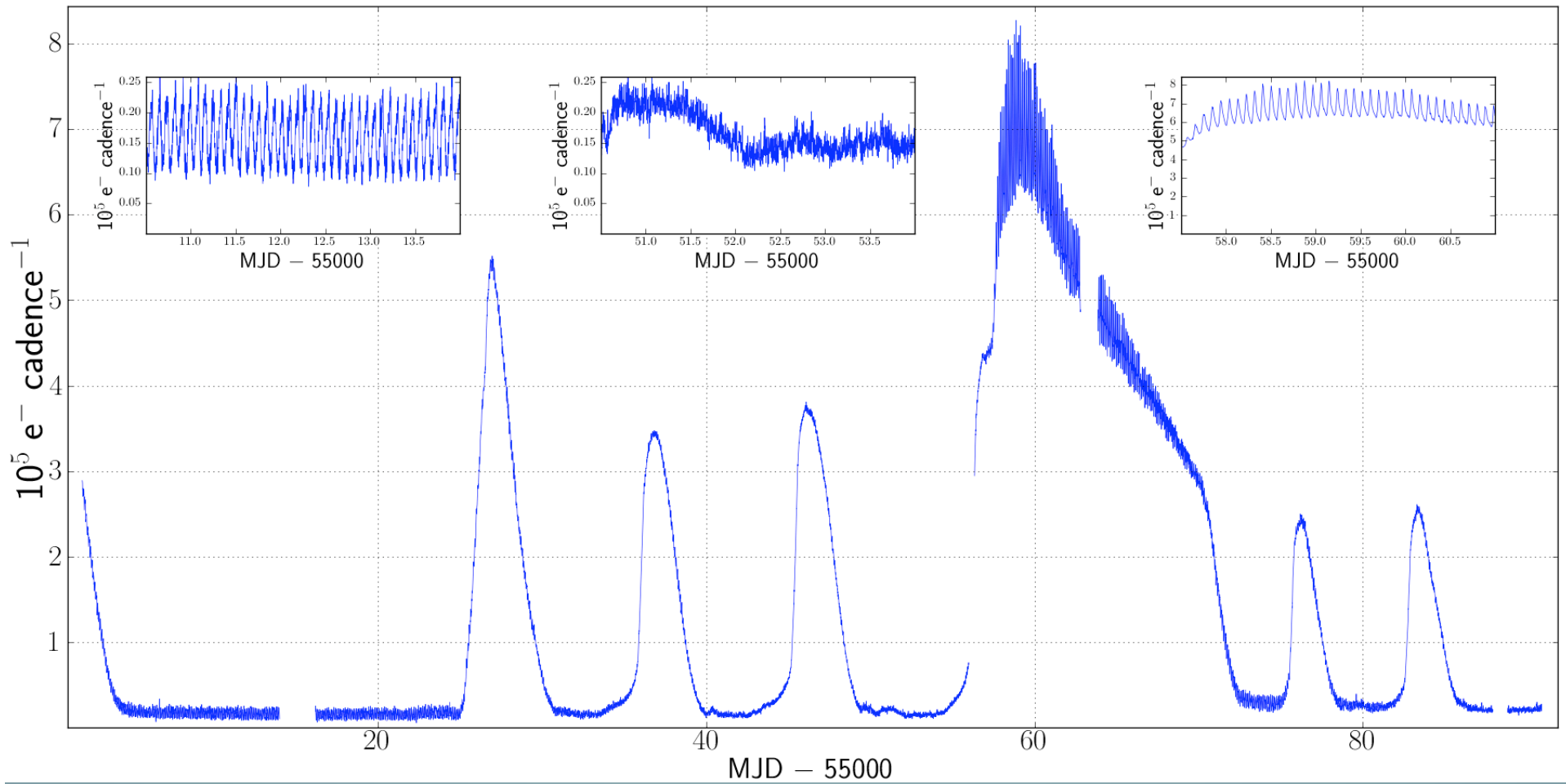


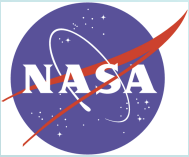


Dwarf Nova – V344 Lyr

Kepler

A Search for Earth-size Planets





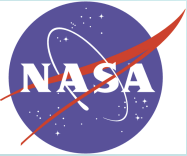
Kepler Planets – Real or Not?

Kepler

A Search for Earth-size Planets

False Transits can be caused by:

1. M dwarfs transiting giants and supergiants
2. White dwarfs transiting solar-type stars
3. Grazing eclipses of one star by another
4. Full eclipses in a faint background binary whose light is blended with a foreground bright star
5. Triple systems – difficult in short term
6. Other even more insidious effects

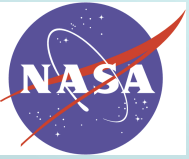


VALIDATION of PLANETS



A Search for Earth-size Planets

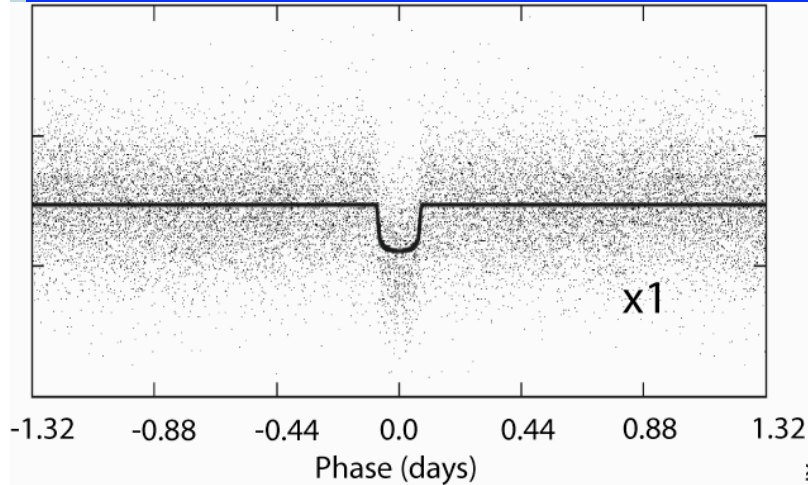
- **SNR > 7, to rule out statistical fluctuations**
- **Three or more transits to confirm orbital periodicity**
- **Light curve depth, shape, and duration consistent**
- **Image subtraction to identify signals from background stars**
- **Radial velocity**
 - **Medium precision to rule out stellar companions**
 - **High precision to measure mass of super-Earths and giant planets**
 - **Rossiter-McLaughlin effect to confirm orbiting planet**
- **High spatial resolution images to identify extremely close background stars; Observe eclipse of background stars.**
- **Check for color change during transit**
- **More tests as the mission progresses and planets get smaller**



Exo-Planet HAT P-7b Kepler Light Curve

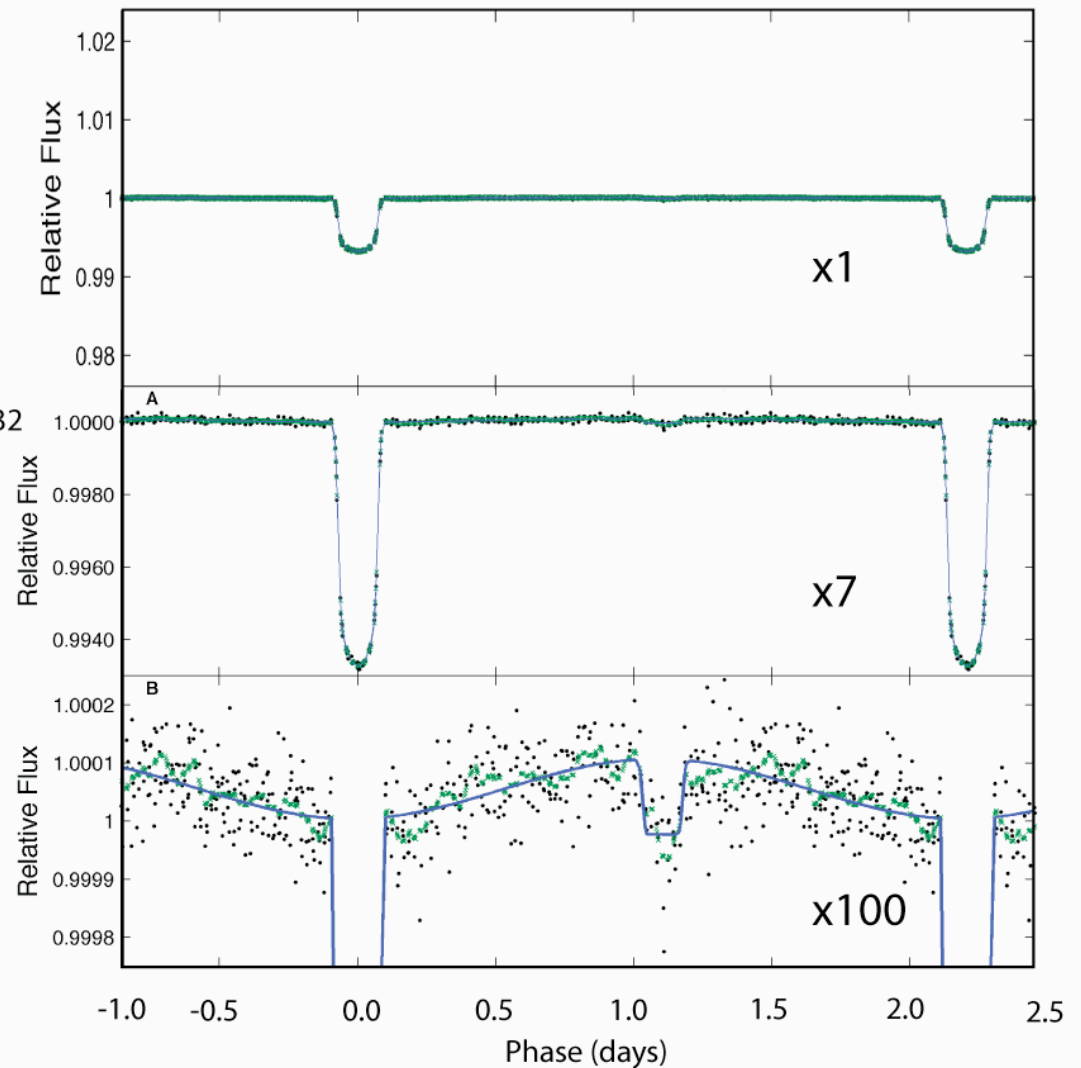


A Search for Earth-size Planets



16,620 HATNet data points (57.7 days of data)

HAT-P-7b data from the ground
A. Pal et al., 2008



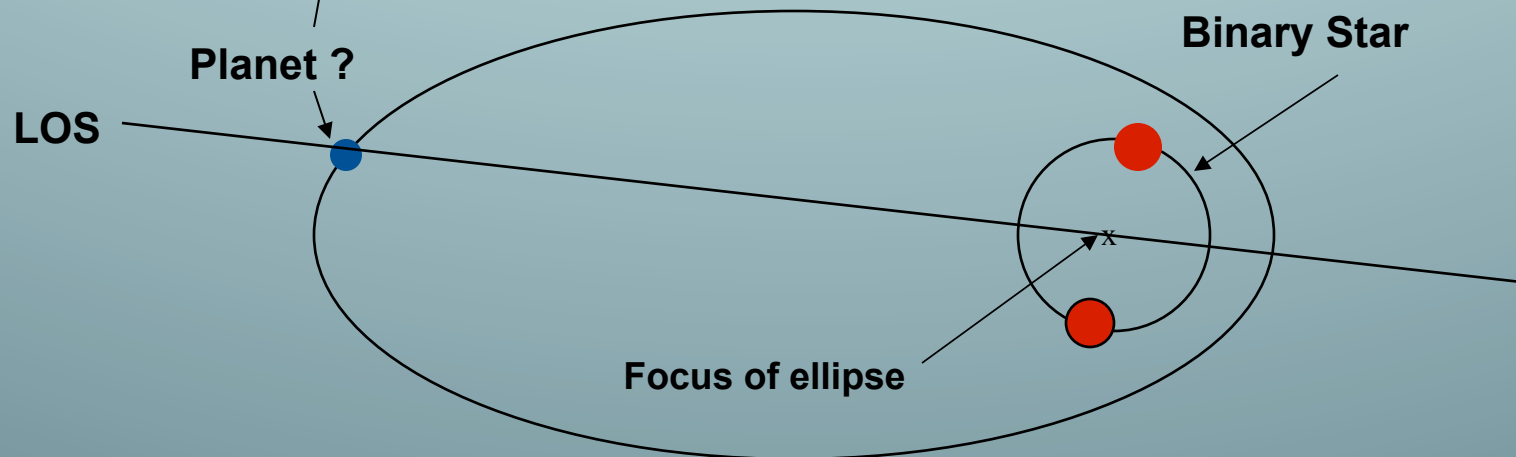
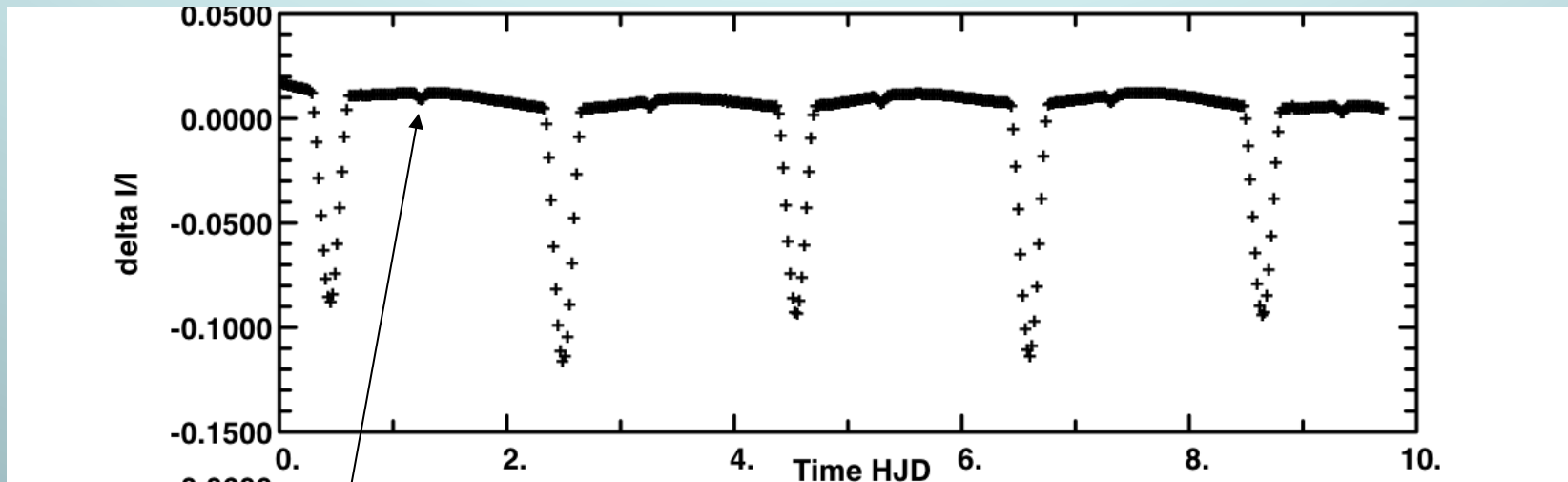
Kepler Commissioning data (10 days)
W. Borucki et al., 2009

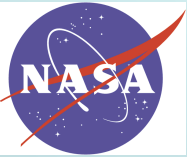


BINARY WITH CIRCUMBINARY PLANET?

Kepler

A Search for Earth-size Planets





NASA HQ View ;-)

Kepler

A Search for Earth-size Planets

“Kepler will answer at least one big question:

Are there other planets like ours in the universe?”

